Interpreting Early Triassic (Smithian) sea-level change and climate using sequence stratigraphy and oxygen isotopes of conodont apatite

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INTERPRETING EARLY TRIASSIC (SMITHIAN) SEA-LEVEL CHANGE AND CLIMATE USING SEQUENCE STRATIGRAPHY AND OXYGEN ISOTOPES OF CONODONT APATITE

BY

STEPHANIE LYNN YURCHYK

B.S. Geology, University of Rochester, 2007

THESIS

Submitted in Partial Fulfillment of the Requirements for the Degree of

Master of Science
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The University of New Mexico
Albuquerque, New Mexico

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UNDERSTANDING EARLY TRIASSIC (SMITHIAN) SEA-LEVEL CHANGE AND CLIMATE USING SEQUENCE STRATIGRAPHY AND OXYGEN ISOTOPES OF CONODONT APATITE

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ABSTRACT OF THESIS

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ABSTRACT

The Early Triassic is conventionally interpreted to have a warm and ice-free climate. During this time, three globally recognized depositional sequences developed in response to ~My-scale sea-level changes. The Lower Triassic Lower Thaynes Formation records the Smithian (2nd sea-level cycle in the Early Triassic) in the western United States (Confusion Range and Weber Canyon, Utah). The Smithian portion of the Mikin Formation records an approximately time-equivalent sea-level cycle in northern India (Guling, Himachal Pradesh). The maximum flooding zone combined with existing age-diagnostic biostratigraphy allows for correlation between two sections in the western United States and one section in northern India suggesting the Smithian My-scale sea-level change was likely eustatic.

Samples were collected for oxygen isotopic analyses of conodont apatite from the two field locations (in Utah) in a sequence stratigraphic framework to better understand the Smithian paleoclimate. Due to an up to ~1.5 ‰ disparity of values between the two locations, additional conodont elements were analyzed from the Guling, Himachal Pradesh, northern India, Bear Lake, southern Idaho, Wapiti Lake area in eastern British Columbia, Canada, and Sverdrup Basin in the Canadian Arctic. Conodont Alteration
Index or CAI (a measure of post-burial thermal alteration based on color) was estimated for each of the locations listed above and range from 1.5 for Wapiti Lake and the Confusion Range to 5 for Guling. In addition, SEM images were taken to identify potential physical alteration of the conodonts for the Confusion Range (smooth surface with no signs of alteration), Weber Canyon (pitted surface with signs of potential alteration), and Guling (visibly the most pitted surface with signs of potential dissolution).

The δ¹⁸O values for Weber Canyon range from ~14.4 to 15.8 ‰, the Confusion Range from ~16 to 16.9 ‰, northern India from ~15.8 to 16.5 ‰, Wapiti Lake from ~17.2 and 17.6 ‰, Sverdrup Basin range from ~14.5 and 14.8 ‰, and the Bear Lake value was ~ 16.5 ‰. Conodonts with a CAI of ≤ 3 or lower produced δ¹⁸O values that most likely reflect the primary Smithian ocean isotopic values. Assuming an ice-free ocean value of -1 ‰, sea-surface temperatures were calculated as ~35 to 38 °C for the paleotropical and ~32 to 34 °C for the paleosubtropical regions, which make sense given their latitudinal position. Warm ocean currents in the neo-Tethys Sea can potentially explain this discrepancy. All six locations indicate that the Smithian ocean was significantly warmer than the present ocean, and instead, most resemble the extreme greenhouse sea-surface temperatures calculated for mid- to late Cretaceous.
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Preface

The aim of this research project is to better understand the interaction between Early Triassic (Smithian) climate and sea-level changes through a detailed sedimentological and sequence stratigraphic perspective (Chapter 1), as well as through the use of δ¹⁸O of conodont apatite as a proxy for sea surface temperature and temperature gradients between the equator and mid paleolatitudes (Chapter 2). This project is important because it adds to the growing amount of detailed stratigraphic work during the recovery period after the Permo-Triassic extinction and evaluates the relatively new tool of using δ¹⁸O of conodont apatite to better understand the Smithian climate.

The purpose of Chapter 1 is to create a sequence stratigraphic framework for the Early Triassic (Smithian), specifically looking at developing a more detailed global correlation. This chapter takes the previous research one step further by combining modern advances in biostratigraphy with detailed sedimentary facies analyses and depositional environment interpretations of the Smithian in the western United States and northern India. This will allow for both a more complete global sequence stratigraphic picture, as well as help to create a framework for understanding the Smithian climate.

In addition to the detailed stratigraphic analysis, samples were collected in a sequence stratigraphic framework for conodont apatite. Chapter 2 will focus on the reliability of δ¹⁸O values from conodont apatite, and the potential for using this tool as a proxy for seawater temperature, ice volume, evaporation, and freshwater influx. Those samples interpreted to retain primary or near primary ocean δ¹⁸O values were used to address the Smithian climate.
To date, it is difficult to identify diagenetic alteration that would allow for the exchange of oxygen between pore waters and conodont elements. This chapter will investigate a number of Smithian conodont elements with Conodont Alteration Indexes (CAIs) ranging from 1.5 (little thermal alteration) to 5 (moderate to extreme thermal alteration), as well as SEM images of the conodonts to look for physical evidence of alteration to better understand why there is an up to ~1.5‰ difference between relatively closely spaced samples collected for this study.
Chapter 1: Sequence Stratigraphy of the Lower Triassic (Smithian) Lower Thaynes Formation (Weber Canyon and Confusion Range, Utah) and the Smithian portion of the Mikin Formation (Guling, Himachal Pradesh, India)

Abstract

The mixed siliciclastic and carbonate Lower Triassic (Smithian) Lower Thaynes Formation at Weber Canyon, northern Utah and the Confusion Range, western Utah are interpreted as ~My-scale transgressive-regressive cycles. They were deposited in an inland sea on the western edge of Pangaea separated from the Panthalassa Ocean by a volcanic arc. The Lower Thaynes Formation conformably overlies the siliciclastic-rich, restricted shallow subtidal to continental Woodside Shale to the north and east and carbonate-rich, shallow marine Dinwoody Formation to the west and south. It is overlain by the Decker Tongue to the north and east and its shallow subtidal to marginal marine red siltstone and shale facies equivalents to the south and west. Facies relationships between the Woodside Shale, Lower Thaynes Formation, and the Decker Tongue and marginal marine equivalents indicate a deepening into open marine facies dominated by bivalves, gastropods, echinoderms, and ammonoids in the transgressive systems tract to deep subtidal, dark shale in the maximum flooding zone, and a shallowing into shallow and restricted shallow subtidal facies in the highstand systems tract. The two sections are correlated both by sequence stratigraphy and existing ammonoid and conodont biostratigraphy.
The carbonate-rich Lower Triassic Smithian portion of the Mikin Formation in Guling, Himachal Pradesh, northwestern India represents the time-equivalent My-scale sea-level change from the southern neo-Tethys Sea on the eastern Margin of Pangaea. The Smithian portion of the Mikin Formation conformably overlies the Griesbachian Lower Limestone Member and the Dienerian portion of the Limestone and Shale Member. The Spathian portion of the Niti (Nodular) Limestone Member of the Mikin Formation overlies the Smithian portion of the Mikin Formation. The Indian section is correlated with the western United States sections using the maximum flooding zone, defined as the deposition that occurs during the maximum transgression. Further support for the correlation comes from ammonoid and conodont biostratigraphy.

The transgressive-regressive sequence for all three sections falls within the bounds of the Smithian time period based on the known biostratigraphy. Thus, it is possible that the sea-level changes are eustatic in origin. However, due to the complexities involving regional changes in accommodation and lack of absolute age dates and the scarcity of age-diagnostic fossils in the Early Triassic, determining the timing of the sea-level change between the two regions is challenging.
1.1 Introduction

The Permian-Triassic extinction (~252 Ma) is recognized as one of the “big five” extinctions in Earth’s history marked by the loss of 84-93% of marine genera (Raup, 1979; Sepkoski, 1989; Stanley, 1990; Erwin, 2006; 2008). The ~5 My recovery period (Early Triassic) is characterized by low biotic diversity (Schubert and Bottjer, 1995; Rodland and Bottjer, 2001; Payne et al., 2006), the absence of corals and sponges (Hallam and Wignall, 1997), the persistence of cosmopolitan marine fauna (Hallam, 1981, Boyer et al., 2004; Fraiser and Bottjer, 2004), continued low deep-water oxygen levels (Wignall and Hallam, 1996), and a hiatus of coal deposition (Retallack et al., 1996). The prolonged recovery is thought to be the result of persistent, unfavorable paleoenvironmental conditions throughout the Early Triassic, though the nature of these paleoenvironmental conditions are not well understood (Erwin, 2006; 2008; Hallam and Wignall, 1997; Payne et al., 2004).

The climate during the Early Triassic is thought to have been warm and stable; however, the assumed stability may be the result of a lack of detailed paleoclimate studies. Evidence to support a globally-warm Early Triassic is based on increased weathering rates in high paleolatitude soils (Retallack, 1999; Michaelsen and Henderson, 2000; Michaelsen, 2002), migration of characteristically warm-water taxa to higher latitudes (i.e., calcareous algae into the Svalbard Basin); (Wignall et al., 1998) and warm weather flora (i.e. the appearance of cycads in the Greenland record); (Looy et al., 2001). The lack of glacial deposits suggests that the poles were not cold enough to sustain permanent ice caps (Frakes et al., 1992). Ammonoid fauna are found at a wide paleo-latitude range through the Smithian suggesting that there is a gentle low to high
latitude temperature gradient (Brayard et al., 2006). General circulation models across the Permo-Triassic boundary (Kidder and Worsley, 2004; Kiehl and Shields, 2005) and geochemical models (Berner et al., 2002; Royer et al., 2004) suggest that there was a 6 to 8 °C increase in the temperature.

Although the Early Triassic is an important and unique period in Earth history, relatively few detailed stratigraphic studies exist, which are necessary to build a framework for global marine sequence correlation. The groundbreaking work identifying genetically related sedimentary packages (or depositional sequences) involved seismic stratigraphy, which lacked relative and absolute time controls (Vail et al., 1977; Haq et al., 1987; 1988) (Figure 1.1a). Embry (1988) combined Triassic outcrop and well-log stratigraphic study from the Sverdrup Basin in the Canadian Arctic to test the Triassic portion of the Phanerozoic eustatic sea-level curve developed by Vail et al. (1977), Haq et al. (1987; 1988). From this study, three sea-level changes were described for the Early Triassic (Figure 1.1b). Additional Early Triassic stratigraphic studies (Embry, 1997; Paull and Paull, 1994; Baud et al., 1996) have broadly identified a similar three Lower Triassic transgressive-regressive sequences: the first in the Griesbachian-Dienerian, the second in the Smithian, and the third in the Spathian. Embry (1997) attempted to correlate the Triassic across seven globally distributed regions; however, this project lacked both adequate biostratigraphic control for these regions and detailed stratigraphy to prove that the sequences were, in fact, correlative.

The purpose of this project is to 1) identify depositional environments within the Lower Triassic Lower Thaynes Formation (western United States) and the Smithian
**Figure 1.1:** a) The Phanerozoic sea-level curve interpreted from both seismic stratigraphy, well logs and outcrops studies, highlighting known known ice ages with blue bars (modified from a compilation of seismic stratigraphy completed by Haq et al., 1988, Ross & Ross 1987, Ross & Ross 1988). b) The above sea level curve highlights the Early and Middle Triassic sea-level curve developed from both seismic and outcrop studies in the Sverdrup Basin in the Canadian Arctic with the Smithian sea-level rise and fall highlighted in the purple box (modified from Embry, 1988).
portion of the Mikin Formation (Guling, Himachal Pradesh, northern India), 2) characterize vertical stacking patterns in each section in order to develop a sequence stratigraphic framework for the Smithian and 3) use published ammonoid and conodont biostratigraphy (Smith, 1969; Garzanti et al., 1994; Krystyn et al., 2004; 2007; Galfetti et al., 2007; Brayard et al., 2009; Guex et al., 2010; Stephen et al., 2010; Beatty, personal communication; Jenks, personal communication) to constrain the age and help correlate the Smithian sequences, which will provide the framework for understanding the origin of the Smithian sea-level change. The Lower Thaynes Formation in Weber Canyon and the Confusion Range and the Smithian portion of the Mikin Formation in Guling, India are chosen for this study because they have well-exposed marine strata with established biostratigraphic age control in geographically diverse locations.

1.2 Geologic Background: Paleogeography and Lower Triassic Stratigraphy

The Pangaean supercontinent and the Panthalassa Ocean (Figure 1.2) dominated the Early Triassic paleogeography. Pangaea remained a single landmass transected by the paleoequator until the Early to Middle Jurassic (Blakey, 2010; Scotese, 2010). Global sea level was relatively low during the late Paleozoic through the early Mesozoic due to low mid-ocean ridge activity (Forney, 1975; Holser and Magaritz, 1987; Ross and Ross, 1987) (Figure 1.1b).
1.2.1 Western North America

During the Early Triassic, western North America lay between ~5 and 35 °N paleolatitude (Figure 1.2) (Blakey, 1974; 2010; Scotese, 2010). This region underwent a complex tectonic history that resulted from the initiation and evolution of two temporally distinct subducted lithospheric slabs: a pre-Farallon plate and the Farallon plate beneath the western edge of North America (Lawton, 1994). Subduction of the pre-Farallon plate was initiated in the Early Triassic following the latest Permian to early Triassic Sonoma Orogeny, which marked the closure of a marginal basin west of North America (Lawton, 1994). Following the brief formation of the earliest Triassic flexural basin (including the Havalleh Basin) adjacent to the Golconda allochthon, initiation of subduction created a north-westward-thickening wedge of marine sedimentary rocks in the marine foreland basin (Smith, 1969; Silberling, 1975). During the Smithian maximum transgression, this inland sea extended from southeastern Utah to British Columbia with an ~160 to 500 km-wide ramp and was separated from the Panthalassa Ocean by a volcanic arc to the west (Blakey, 1974).

Subsidence within the foreland basin is attributed to the viscous flow of the mantle above the slab of the subducted oceanic crust, tectonic loading by the thrust sheets and crustal loading by sediments following the closure of a backarc basin during the Sonoma Orogeny (Lawton, 1994). Significant orogenic features that influenced siliciclastic sedimentation in the Havalleh basin included the Uncompahgre Highland, the Kaibab Uplift, the Ancestral Rocky Mountains, and the Mogollon Highland to the east and south (Blakey, 1974; Paull and Paull, 1994). Fluvial sediment transport into the basin occurred from the east and southeast (Marzolf, 1994).
Figure 1.2: Early Triassic paleogeographic map. The supercontinent of Pangaea is fully assembled with the Panthalassa Ocean surrounding the eastern margin and the paleo- and neo-Tethys Seas bordering the eastern margin, dotted with microcontinents and volcanic arcs. Exposed land masses are dark gray and submerged continental platforms are light gray. Sites 1 and 2 are Weber Canyon, northern Utah and Confusion Range, western Utah, respectively, and were deposited in an intracratal basin separated from the open Panthalassa Ocean by a volcanic arc. Site 3 is Guling, northern India deposited on a passive margin that dipped to the north towards the neo-Tethys Sea (modified from Blakey, 2010 and Scotese, 2010).
1.2.2 The Thaynes Formation (Smithian and Spathian)

The Thaynes Formation (~300 to 1100 m thick) was originally described by Boutwell (1907) in Park City, Utah and ranges from deep subtidal to marginal marine deposits that commonly consist of skeletal limestone, calcareous siltstone and shale facies. The Thaynes Formation overlies the Griesbachian-Dienerian shallow marine to continental Woodside Shale in northern and eastern Utah and the shallow-water carbonates of the Dinwoody Formation to the north and west in Idaho, western Utah, Nevada, and Montana (Kummel, 1943; 1954) (Figure 1.3). The Woodside Shale and Dinwoody Formation unconformably overlie the Upper Permian Phosphoria Formation in most of the basin; however, in western Utah and Nevada, the Thaynes Formation unconformably overlies the Upper Permian Gerster Formation (Kummel, 1954; Collinson, 1976).

The Thaynes Formation is separated into the Lower Thaynes (Smithian) and the Upper Thaynes (Spathian) Formations divided by the Decker Tongue to the east and very shallow marine deposits to the west (Kummel, 1954; Smith, 1969) (Figure 1.3). The Ankareh Formation of the Moenkopi Group is characterized by continental sandstone and shale and overlies the Thaynes Formation (Kummel, 1954). Regional correlation between the Thaynes Formation in Utah, Idaho, and Nevada were originally defined using ammonoid biostratigraphy (early Smithian Meekoceras ammonoid zone (Boutwell, 1912). Lucas et al. (2007) argued that some of the ammonoids that were originally defined as Meekoceras were, in fact, misidentified due to poor preservation. Instead, the ammonoid zone used for regional correlation of the maximum transgression is the mid- to late Smithian Anasibirities ammonoid zone (Lucas et al., 2007).
Figure 1.3: Chronostratigraphic diagram for the Lower Triassic in the western United States (on the left; modified after Lucas et al., 2007) and northern India (on the right; modified from Krystyn et al., 2007). Locations discussed in this study are in bold. Smithian ammonoid biostratigraphy is also listed for the western United States (modified from Kummel and Steele, 1962), as well as for the northern India section (modified from Krystyn et al., 2007). Absolute ages of ~252 Ma for the Permo-Triassic boundary and ~247 Ma for the Early and Middle Triassic boundary are included (Ovtcharova et al., 2006). These ages are from recorded using U-Pb dating techniques on ash beds from northwestern Guangxi, South China and correlated globally using ammonoid biostratigraphy (Ovtcharova et al., 2006).
More recently, advancements in ammonoid and conodont biostratigraphy have made it possible to build a framework for global correlation. Using the absolute U/Pb ages from the time-equivalent Jinya section in China combined with ammonoid and conodont biostratigraphy, the Griesbachian through Smithian is \( \sim 2 \text{ My} \) in duration and the Spathian as \( \sim 3 \text{ My} \) (Ovtcharova et al., 2006). This suggests that the entire Smithian represents \( \leq 2 \text{ My} \).

The mixed carbonates and siliciclastics of the Lower Thaynes Formation at Weber Canyon (\( \sim 240 \text{ m} \) thick) was measured and described on the north and south sides of the Weber River, adjacent to the Union Pacific Railroad tracks and west of Devils Slide in Morgan County, Utah (Exit 108 off of I-84E; GPS location: 12 T 0452004 UTM 4545695) (Figure 1.4; Appendix A). Weber Canyon is the best-known exposure of the Lower Thaynes Formation in western North America (Figure 1.4). Most outcrops of the Thaynes Formation only expose the base and the top of the formation because the middle portion is composed of fine-grained silty lime mudstone and mudshale facies, which are often covered. At this locality, the marine Lower Thaynes Formation conformably overlies red-brown sandstone, siltstone, and shale of the Woodside Shale (Griesbachian-Dienerian) and is conformably overlain by a similar red-brown marginal marine to continental siliciclastic unit (Decker Tongue) at Weber Canyon (Kummel, 1954; Smith, 1969). The lower contact is defined by the first and the upper contact by the last beds containing marine macrofossils (Smith, 1969).

At Weber Canyon, the Lower Thaynes Formation consists of three depositional units: 1) the carbonate-rich lower unit composed primarily of gray, silty lime mudstone containing sparse to rare fossils and wackestones to grainstones containing bivalves,
Figure 1.4: The Lower Triassic Lower and Upper Thaynes Formations exposed at Weber Canyon along I-84 east of Ogden, Utah. The contact between the silty lime mudstone (top of unit 1) and the dark calcareous shale (unit 2) is labeled (~153 m). Also, the top of the Lower Thaynes Formation (~240 m; top of unit 3) is labeled. The Decker Tongue conformably overlies the Lower Thaynes Formation and the Spathian Upper Thaynes Formation, which conformably overlies the Decker Tongue, are also labeled.
echinoderms, and gastropods (~150 m), 2) a dark shale devoid of fossils that shallows into a calcareous mudstone interbedded with thin carbonates (~40 m), and 3) an upper mixed siliciclastic and carbonate unit of silty lime mudstone and calcareous siltstone, interbedded with whole bivalve wackestones (~70 m). Age-diagnostic fossils in the Lower Thaynes Formation, though scarce, consists primarily of upper Smith ammonoids \textit{(Anasibirities)} and mid- to upper Smithian conodonts \textit{(N. waageni)} (Beatty, T., written communication; Jenks, J., personal communication; Smith, 1969). The first beds containing early Spathian conodonts are stratigraphically above the Decker Tongue and help constrain the marine Lower Thaynes Formation as likely Smithian in age.

The Lower Thaynes Formation in the Confusion Range (~250 m) was measured and described at two locations: Disappointment Hills (GPS location: 12S 0268549 UTM 4365751) and Cowboy Pass, separated by ~11 kilometers (Figure 1.5). In the Confusion Range, four depositional units comprise the Lower Thaynes Formation: 1) a resistant, light olive-gray and pinkish wackestone to grainstone (~10 m), 2) primarily light olive-gray to yellowish calcareous mudshale occasionally capped by ammonoid and micro-gastropod rich limestone beds (.25 to 2 m thick) (~70 m), 3) a light yellow mudshale (~100 m) (Hose and Repenning, 1959), and 4) a red mudshale to fine sandstone interbedded with bivalve-rich limestone beds (~70 m) (Collinson, 1979). Due to the post-depositional complex tectonic history at Cowboy Pass, only the first and fourth
Figure 1.5: The Lower Triassic Lower Thaynes Formation exposure in western Utah (Confusion Range, Disappointment Hills). Units 1 (~10 m) and 2 (~70 m) are labeled. Here the Lower Thaynes Formation overlies the Permian Gerster Formation.
depositional units are well exposed, whereas at Disappointment Hills only the first and second units can be examined in a continuous section. Unit 3, a yellow mudshale, does not outcrop well at either location, but is estimated to have a thickness of ~100 m (Hose and Repenning, 1959). Despite having to piece together the Confusion Range, these sections were chosen for this study due to the abundance of age-diagnostic fossils.

The Lower Thaynes Formation in the Confusion Range unconformably overlies the Permian Gerster Formation, which is overlain by a conglomerate varying in thickness from 0 to 1.5 m. This unconformity extends from western Utah to eastern Nevada and has been recorded by several authors (Newell, 1948; Wheeler et al., 1949; Clark, 1956; 1959; Bissell, 1962a, 1962b; Collinson, 1968). Collinson et al. (1976) investigated the origin of this unconformity and concluded that it resulted in significant time loss likely caused by uplift and erosion related to the Sonoma Orogeny. Newell (1973) suggested that a eustatic drop in sea level caused the hiatus. Shallow marine red-brown sandstones and siltstones conformably overlie the Lower Thaynes Formation (~50 m; Atudorei, V., personal communication).

Ammonoid fossils representing the *Meekoceras, Inyoites* and *Anasibirities* zones are abundant near the base of the Lower Thaynes Formation at the Confusion Range (Brayard et al., 2009; Stephen et al., 2010). At the top of the Lower Thaynes Formation, however, no age diagnostic marine fossils have been recovered (Hose and Repenning, 1959; Newell, 1959). Thus, the only marker to distinguish the Lower from the Upper Thaynes is the first Spathian ammonoid-bearing bed in the field area (*Tirolites*) (Guex et al., 2010). The *Tirolites* lies above the red sandstone and shale facies (Guex et al., 2010).
1.3 Facies Descriptions and Depositional Environment Interpretations

Two Lower Thaynes Formation from sections from Weber Canyon and the Confusion Range were measured and described on a bed-by-bed basis for facies descriptions and paleoenvironmental analysis (Table 1; Appendix A). Facies and internal subfacies were originally defined in the field based on grain size, bedding characteristics, sedimentary structures, biota, and facies associations; 27 thin sections were utilized to refine the facies descriptions and interpretations (Tables 1.1, 1.2, 1.3). Facies stacking patterns were used to define ~My-scale depositional sequences. Interpretations of paleoenvironmental change and sea-level change were then derived from facies stacking patterns. The inner- to outer-ramp terminology, defined by Burchette and Wright (1992), is used for depositional environment interpretations. Vertical facies relationships of the cycles will be used to describe changes in ramp morphology through time.

1.3.1 Weber Canyon, northern Utah

At Weber Canyon, four depositional facies and seven subfacies are recognized, including: restricted shallow subtidal (silty lime mudstone, abraded bivalve wackestone-packstone), shallow subtidal (echinoderm grainstone), intermediate subtidal (bivalve wackestone), and deep subtidal (silty skeletal lime mudstone, lime mudstone interbedded with silty dark mudshale interbeds (rhythmites), and dark calcareous mudshale) (Table 1.1).
<table>
<thead>
<tr>
<th>Subfacies</th>
<th>Bedding</th>
<th>Skeletal components</th>
<th>Sedimentary-biologic features</th>
<th>Environment of deposition</th>
</tr>
</thead>
<tbody>
<tr>
<td>Silty lime mudstone</td>
<td>Thin-medium</td>
<td>none observed</td>
<td>suspension laminations, sparse symmetric ripples</td>
<td>Restricted shallow subtidal, low to moderate energy, inner ramp</td>
</tr>
<tr>
<td>Abraded bivalve wackestone-packstone</td>
<td>Medium-thick</td>
<td>bivalves, sparse echinoderms, micro-gastropods, (conodonts)</td>
<td>bioturbation, occasional cross bedding</td>
<td>Restricted subtidal, moderate energy, inner ramp</td>
</tr>
<tr>
<td>Echinoderm grainstone</td>
<td>Medium</td>
<td>echinoderms, abraded bivalves, (conodonts), sparse abraded micro-gastropods</td>
<td>cross bedding</td>
<td>Shallow subtidal, high energy, inner to mid ramp</td>
</tr>
<tr>
<td>Bivalve wackestone</td>
<td>Medium</td>
<td>bivalves, sparse to rare micro-gastropods, and echinoderms, (conodonts)</td>
<td>bioturbation</td>
<td>Intermediate subtidal, low energy, inner to mid ramp</td>
</tr>
<tr>
<td>Silty skeletal lime mudstone</td>
<td>Thin-medium</td>
<td>sparse-rare abraded bivalves, micro-gastropods, echinoderms (conodonts)</td>
<td>locally disrupted laminae (bioturbation), strongly laminated toward the top, sparse symmetric ripples and thin ripples with clay drapes</td>
<td>Deep subtidal, low energy, periodically poorly oxygenated (at times), outer ramp</td>
</tr>
<tr>
<td>Lime mudstone interbedded with silty dark mudshale (rhythmites)</td>
<td>Thin bedding</td>
<td>rare abraded bivalves (conodonts) only in the lime mudstone</td>
<td>lime mudstone is bioturbated</td>
<td>Deeper Subtidal, low energy, poorly oxygenated, outer ramp</td>
</tr>
<tr>
<td>Dark calcareous mudshale</td>
<td>Thin</td>
<td>none observed</td>
<td>none observed</td>
<td>Deepest subtidal, low energy, outer ramp</td>
</tr>
</tbody>
</table>
**Restricted shallow subtidal facies**

*Silty lime mudstone subfacies:*

The silty lime mudstone subfacies ranges in thickness from <1 to 30 cm and is deposited near the base and top of the Lower Thaynes Formation (Figure 1.6). This subfacies was deposited in moderate energy, poorly oxygenated shallow subtidal, inner-ramp environments. This interpretation is supported by a lack of subaerial exposure features, fine-grain size, suspension laminations, sparse symmetric ripples, and lack of both fossils and bioturbation. In addition, this subfacies is in a conformable stratigraphic position above the interpreted continental beds of the Woodside Shale and below open-marine subfacies. The abundant, well-sorted quartz silt grains indicate a consistent current energy and/or eolian source and the fine-grain size indicates that this subfacies was deposited distal from its source.

*Abraded bivalve wackestone-packstone subfacies:*

The abraded bivalve wackestone-packstone subfacies ranges in thickness from 0.1 to 1 m and was deposited in moderate energy, shallow subtidal open marine, inner-ramp environments (Figure 1.7). The bioclasts are abraded, disarticulated, and moderately to well sorted, which suggests that they were extensively reworked in and transported from higher energy environments. The abundance of lime mud suggests moderate energy environments. The bivalves and gastropods likely filled an open marine, shallow subtidal niche that was vacated by organisms, such as brachiopods, during the time of ecosystem recovery following the end-Permian extinction (Boyer et al., 2004). Sparse echinoderm grains further suggest occasional open-marine circulation or onshore transport of offshore
Figure 1.6: Field photograph of thin and medium bedded silty lime-mudstone subfacies. Field book for scale.
Figure 1.7: Field photograph of the abraded bivalve wackestone-packstone outcrop. 
Rock hammer for scale.
crinoid bioclasts (Lehrmann et al., 2001). Based on its stratigraphic association with inner-ramp silty lime mudstone, the wackestone-packstones are interpreted to have been deposited during times of increased energy when either waves were able to surpass the shoal barrier (echinoderm grainstone subfacies) and/or when the shoal barrier was absent.

**Shallow subtidal facies**

*Echinoderm grainstone subfacies:*

The echinoderm grainstone subfacies ranges in thickness from 0.5 to 1 m and was deposited in open marine, high energy, mid-ramp environments. This interpretation is based on the relatively high faunal abundance and diversity, cross bedding, and lack of mud. In addition, its stratigraphic association with both fine-grained, low-energy inner- and outer-ramp suggests that it may have served as a barrier and sediment source for the lower energy deposition (i.e. abraded bivalve wackestone-packstone and the silty skeletal lime mudstone). The well-sorted and abraded nature of the bioclasts indicates significant transport and, thus, the organisms were likely not living in this environment.

**Intermediate subtidal facies**

*Bivalve wackestone subfacies:*

The bivalve wackestone subfacies ranges in thickness from 0.1 to 0.5 m and was deposited in low energy, inner- to mid-ramp environments indicated by the presence of bioturbation and stratigraphic association with both deep subtidal (rhythmites subfacies) and restricted shallow subtidal (silty lime mudstone) deposits (Figure 1.8). The whole, unabraded shells indicate that the bivalves were not extensively transported or reworked.
Figure 1.8: Field and hand sample photograph of the bivalve wackestone subfacies. The facies is lighter in color and the calcareous mudshale is interbedded and dark in color. Gallon-size plastic bag for scale.
and represent local communities living in or very close to this environment. Boyer et al. (2004) reports similar bivalve-dominated facies in Lower Triassic rocks as monospecific shell beds and interpreted that they were deposited in a low to moderate energy, mixed carbonate-siliciclastic mid-ramp environments.

Deep subtidal facies

*Silty skeletal lime mudstone subfacies:*

The ~90 m-thick silty skeletal lime mudstone subfacies, located in the middle of the Lower Thaynes Formation (~50% of unit 1), represents the most abundant subfacies and was deposited in low energy, poorly to moderately oxygenated outer-ramp environments. Evidence to support this interpretation include the fine-grain size, sparse to rare fossils, clay draped symmetric ripples at the base and horizontal laminations near the top. Terrigenous silt was likely sourced from the exposed craton to the east and south via rivers associated with the Moenkopi Formation and/or eolian sources. This interpretation is further supported by its association with shallower, high-energy deposits at the base and the deepest water subfacies at the top indicating a deepening upward trend. The intensity of bioturbation varies throughout indicating variable dissolved oxygen contents during deposition. Smith (1977) interpreted a similar dark lime mudstone facies as being deposited in deep subtidal environments in the Mississippian Lodgepole Formation.
Lime mudstone interbedded with dark silty mudshale subfacies (rhythmites):

This subfacies ranges in thickness from 0.5 to 30 m with individual beds from <1 cm to ~5 cm. It was deposited in low energy, poorly oxygenated, outer-ramp environments (Figure 1.9). This interpretation is based upon the fine grain size, lack of current-generated sedimentary structures and bioturbation, sparse to rare fossil content, and its association with the deepest water subfacies at the base. The rhythmic interbedding of 1-10 cm thick lime mudstone and mudshale was likely the result of periodic fluctuations in detrital carbonate influx and/or changes in siliciclastic fluvial influx into the marine basin (Elrick et al., 1991; Elrick and Hinnov, 1996; Elrick and Hinnov, 2007). There is an increase in number and thickness of lime mudstone interbeds up section, concurrent with an increase in fossil content and bioturbation corresponding with a decrease in number and thickness of calcareous mudshale, suggesting an overall upward-shallowing trend.

Dark calcareous mudshale subfacies:

This ~30 m thick, platy subfacies was deposited in low energy, poorly oxygenated outer ramp environments (Figure 1.10). This interpretation is based upon its fine-grain size, lack of current-generated sedimentary structures, and lack of bioturbation and fossils. This subfacies is deposited in an upward-deepening portion of the sequence and the two storm beds at the base (skeletal packstones) indicate that initial deposition took place near storm wave base. The absence of storm beds above the contact between this facies and the shallower water facies below indicates that deposition continued below storm wave base for the majority of time. Toward the top of the subfacies, the shale
Figure 1.9: Field photograph of the lime mudstone interbedded with silty dark mudshale subfacies (rhythmites). Field pack and gallon-sized plastic bag to the right for scale.
Figure 1.10: Field photographs of the dark calcareous mudshale subfacies. There is a sharp contact between the light-colored, silty skeletal lime mudstone subfacies and the dark mudshale.
becomes less fissile, indicating an increase in bioturbation. Dark shale facies can form in a variety of depositional regimes; however, based on the stratigraphic position of the subfacies, it is best interpreted as an anoxic to dysoxic, low energy, outer-ramp environments (Arthur and Sageman, 1994).

1.3.2 Confusion Range, western Utah

In the Confusion Range, four depositional facies and eight subfacies were also identified, with two subfacies similar to those found in Weber Canyon (Table 1.2). In descending water depths they include: restricted shallow subtidal (red silty lime mudstone, vuggy (subtidal microbialite) lime mudstone, and bivalve wackestone), shallow subtidal (clast-supported chert-quartz pebble conglomerate, and micro-gastropod-skeletal packstone-grainstone), intermediate subtidal (ammonoid-micro-gastropod wackestone), and deep subtidal (lime mudstone interbedded with calcareous mudshale (“rhythmites”), and light yellow calcareous clayshale) (Table 2; Appendix A).
<table>
<thead>
<tr>
<th>Confusion Range</th>
<th>Lower Thaynes Formation</th>
<th>Sedimentary and biologic features</th>
<th>Environment of deposition</th>
</tr>
</thead>
<tbody>
<tr>
<td>Subfacies</td>
<td>Bedding</td>
<td>Skeletal Components</td>
<td></td>
</tr>
<tr>
<td>Red silty-lime mudstone</td>
<td>Thin-medium</td>
<td>none observed</td>
<td>Laminations, symmetric ripples, rare cross bedding</td>
</tr>
<tr>
<td>Vuggy (subtidal microbialite) lime mudstone</td>
<td>Thin-medium</td>
<td>none observed</td>
<td>vugs</td>
</tr>
<tr>
<td>Bivalve wackestone</td>
<td>Thin-medium</td>
<td>whole bivalves, micro-gastropods</td>
<td>bioturbation</td>
</tr>
<tr>
<td>Clast-supported chert-quartz pebble conglomerate</td>
<td>Medium-thick</td>
<td>Sparse bivalves and micro-gastropod fragments, chert</td>
<td>poorly sorted, angular clasts</td>
</tr>
<tr>
<td>Micro-gastropod skeletal packstone-grainstone</td>
<td>Medium-thick</td>
<td>Abraded bivalves and crinoids, micro-gastropods, ammonoid fragments</td>
<td>massive, cross bedding, graded beds</td>
</tr>
<tr>
<td>Ammonoid/micro-gastropod wackestone</td>
<td>Thin-medium</td>
<td>Ammonoids, micro-gastropods, bivalves</td>
<td>Bioturbation</td>
</tr>
<tr>
<td>Lime mudstone-calcareous mud-shale rhythmsites</td>
<td>Thin bed</td>
<td>none observed</td>
<td>Bioturbation in mudstone</td>
</tr>
<tr>
<td>Light (yellow) calcareous clayshale</td>
<td>thin</td>
<td>none observed</td>
<td>none observed</td>
</tr>
</tbody>
</table>
Restricted shallow subtidal facies

Red silty lime mudstone:

The red silty lime mudstone subfacies ranges in thickness from <1 cm to 10 cm and is only found at the top of the formation. This subfacies was deposited in low energy, low oxygen, inner-ramp environments. This interpretation is based on lack of both fossil and bioturbation, and the occurrence of suspension laminations, sparse symmetric ripples. The depositional environment of this facies is up for debate (McKee, 1949; Picard, 1967; McCormic and Picard, 1969; Blakey, 1974), but has been interpreted as peritidal to tidal flat environments based on evidence of periodic submersion and its relationship with shallow water carbonate deposits (bivalve wackestone subfacies).

Vuggy (subtidal microbialite) lime mudstone:

The vuggy (subtidal microbialite) lime mudstone subfacies (~0.25 m thick) was deposited in low energy, inner-ramp, shallow subtidal-intertidal environments (Figure 1.11). This interpretation is based on fine-grain size, primary vugs, and stratigraphic relationship with other restricted shallow subtidal subfacies.

Bivalve wackestone:

The bivalve wackestone subfacies ranges in thickness from 0.2 to 1 m and was deposited in low energy, inner- to mid-ramp restricted subtidal environments. This interpretation is based on the presence of bioturbation and its stratigraphic
Figure 1.11: Field photograph of the vuggy (subtidal microbialite) lime mudstone subfacies. This subfacies is only present near the base of the Confusion Range Lower Thaynes Formation. Field book for scale.
relationship with other restricted shallow subtidal subfacies. Boyer et al. (2004) suggests that similar facies with monospecific taxa from other Lower Triassic deposits in western North America indicate a period of recovery. The whole, unabraded nature of the bivalves demonstrates that they were not extensively transported or reworked.

**Shallow subtidal facies**

*Clast-supported chert-quartz pebble conglomerate:*

The clast-supported chert conglomerate subfacies vary laterally in thickness from ~0-0.3 m at Cowboy Pass to ~0-1.5 m at Disappointment Hills (Collinson, 1976) and are interpreted to represents conglomerates shed off a local fault scarp related to the Sonoma Orogeny (Collinson, 1976) or transgressive lag conglomerates deposited during sea-level rise (Newell, 1973). These high energy events took place in the inner-ramp, shallow subtidal environments evidenced by poor sorting of the clasts indicating changes in depositional energies, marine fossil content in the matrix, lack of mud, and stratigraphic position within other shallow subtidal subfacies.

*Micro-gastropod-skeletal packstone-grainstone:*

The micro-gastropod-skeletal packstone-grainstone subfacies is ~6 m thick and was deposited in high energy, well-oxygenated, inner- to mid-ramp environments. This interpretation is based on poor sorting, cross bedding, abundant, abraded fossil content, and its stratigraphic relationship with both restricted shallow subtidal and deep subtidal facies. Like the echinoderm grainstone subfacies from the Weber Canyon section, this subfacies was deposited in a high energy environment that resulted in a sand bar or shoal.
Fraiser and Bottjer (2004) report that these gastropod species (ecological opportunists) could have occupied any of the ramp environments in other Lower Triassic stratigraphic units and, thus could have been transported here from inner to mid-ramp environments. The micro-gastropod-skeletal packstone subfacies is also documented in the time-equivalent, inner-ramp Sinbad Formation in eastern and southeastern Utah (Davidson, 1967; Fraiser and Bottjer, 2004; and Lucas et al., 2007) further suggesting that the fossils were likely transported from a shallower environment where they were living during times of maximum transgression.

Intermediate subtidal facies

*Ammonoid-micro-gastropod wackestone:*

The ammonoid-micro-gastropod wackestone subfacies (Figure 1.12) ranges in thickness from 0.1 to 0.5 m and was deposited in low-energy, mid- to outer-ramp environments. This interpretation is based on abundance of mud and sharp stratigraphic contacts with deep subtidal subfacies. The abraded nature of the ammonoids and gastropods suggests that they were not likely living in this environment and were instead transported here from the inner- to mid-ramp environments where they were likely living. In addition, Stephen et al. (2010) documented that the ammonoid shell sizes range from <1 cm to >30 cm; however, ammonitellas (embryonic and early juvenile phases) were not found suggesting that the ammonoids may have reproduced elsewhere. The ammonoid species documented here are *Meekoceras, Inyoites, and Anasibirities*, which span the entire Smithian (Brayard et al., 2009; Stephen et al., 2010)
Figure 1.12: Field photograph of the maroon-colored, more resistant ammonoid-micro-gastropod wackestone subfacies, which occasionally caps the buff- to yellow-colored lime mudstone interbedded with light calcareous mshale subfacies. Spray paint container for scale.
Deep subtidal facies

*Lime mudstone interbedded with light calcareous mudshale (rhythmites)*:

The lime mudstone interbedded with light calcareous mudshale subfacies (Figure 1.12) ranges in total thickness of ~0.5 to 15 m and represents the majority of the transgressive facies at the Confusion Range. It was deposited in low energy, poorly oxygenated, outer-ramp environments based on the fine-grain size, lack of current-generated sedimentary structures and bioturbation, rare to absent fossils, and its association with the deepest water subfacies. The meter-thick rhythmite intervals are occasionally capped by the Ammonoid-micro-gastropod wackestone-packstone subfacies (0.1 to 0.5 m thick). Similar to the Weber Canyon rhythmites subfacies, the rhythmic interbedding of 1-10 cm thick lime mudstone and mudshale was likely the result of periodic fluctuations in detrital carbonate influx and/or changes in siliciclastic fluvial influx into the marine basin (Elrick et al., 1991; Elrick and Hinno, 1996; Elrick and Hinno, 2007).

*Light (yellow) calcareous clayshale*:

The poorly exposed, light yellow calcareous clayshale subfacies is ~100 m thick, (Collinson, 1976; Hose and Repenning, 1959) and was deposited in low-energy, poorly oxygenated, outer-ramp environments. This interpretation is based on its fine-grain size, and lack of both bioturbation and fossils.
1.3.3 Guling, Himachal Pradesh, northern India

The main focus of the study was to understand the depositional facies relationships in the western United States; however, I also had the opportunity to participate in a research expedition to the Spiti Valley, Himachal Pradesh, northern India (Himalayas) and observe several well-exposed and biostratigraphically well-constrained Lower Triassic sections (Figure 1.13). In the Early Triassic, Northern India sat at ~40 °S paleolatitude in the southern neo-Tethys Sea (Ogg and von Rad 1994). The neo-Tethys Sea, which began to open in the Permian, contained many micro-continents and volcanic arcs and was surrounded on three sides by continents (Baud et al., 1984) (Figure 1.2).

The Mikin Formation was deposited on a passive margin in the southern neo-Tethys Sea during the Early Triassic (Baud et al., 1984). Subsidence of the passive margin in the mid- to late Triassic is likely related to the extensional tectonics that were involved in opening the neo-Tethys Sea (Baud et al., 1984). Uniformity within the Lilang Supergroup in the northern Indian Himalayas (Spiti Valley) suggests that the sediments were deposited on a homogenous epicontinental ramp with little variations in subsidence (Krystyn et al., 2004).

The marine Triassic in the Indian subcontinent is confined to the Himalayas and the Triassic sequence in the Spiti Valley, Himachal Pradesh is considered to have well developed biostratigraphy and stratigraphic continuity (Bhargava et al., 2004; Krystyn et al., 2004; 2007). Originally described by Hayden (1904), with recent updates in the
Figure 1.13: The Lower Triassic Smithian portion of the Mikin Formation in the town of Guling in Himachal Pradesh, northern India. It is bounded by a fault at the base and the Spathian Mikin Formation at the top. Packs for scale.
stratigraphy by Krystyn et al. (2004; 2007) puts the entire Lower Triassic within the Lilang Supergroup (up to 1250 m thick), Tamba Kurkur Group, Mikin Formation (~35 m) (Bhargava et al., 2004). The Mikin Formation is divided into four members: the Lower Limestone Member, the Limestone and Shale Member, the Niti (Nodular) Limestone, and the Himalayan Muschelkalk and is generally composed of concretionary, nodular limestone, cherty argillaceous, sporadic fine shale partings (Bhargava et al., 2004).

1.3.4 The Lilang Supergroup: Mikin Formation (Lower Triassic)

The Lilang Supergroup discomformably overlies the Upper Permian Gungri Formation and conformably underlies the Lower Jurassic Kioto Group Tagling Formation (Bhargava, 2008). The Lilang Supergroup is subdivided into four groups and ten formations based on detailed ammonoid and, when available, conodont biostratigraphy (see Bhargava et al., 2004 and Krystyn et al., 2004; 2005 for full description). Of the ~1250 m of stratigraphy, the Lower Triassic and half of the Middle Triassic (Tamba Kurkur Group Mikin Formation) occupy ~35 m (Bhargava et al., 2004). The Lower Triassic Mikin Formation is subdivided into the Griesbachian Lower Limestone Member, the Dienerian and Smithian Limestone and Shale Member, and the Spathian Nodular (Niti) Limestone Member.

This study focuses on the Smithian portion of the Mikin Formation (~14 m thick) measured and described in the town of Guling (Appendix A). Five ammonoid zones, *Rohilites, Flemingites flemingianus, Meekoceras gracilitatis*, and *Wasatchites- Anasibirities, and Xenocelites*, and one conodont zone, *N. waageni*, define the Smithian
and are used for global sequence correlation (Carr and Paull, 1983; Carr, 1983; and Krystyn et al., 2007). Combining sequence stratigraphic methods with recent advances in ammonoid and conodont biostratigraphy (Krystyn et al., 2004; 2007; Galfetti et al., 2007, Brayard et al., 2009; Stephen et al., 2010) allows for the correlation of the western United States with the northern India Smithian sequence (Figure 1.15).

Smithian portion of the Mikin Formation at Guling, Himachal Pradesh

In Guling, Himachal Pradesh, northern India two facies (intermediate and deep subtidal) and six subfacies (nodular skeletal wackestone, skeletal wackestone, skeletal mudstone, skeletal mudstone-wackestone with mudshale interbeds, dark mudshale interbedded with skeletal packstones, and dark mudshale) are identified (Table 1.3; Figure 1.14).

The two intermediate subtidal subfacies are the nodular skeletal wackestone and the skeletal wackestone subfacies. The nodular skeletal wackestone subfacies ranges in thickness from ~0.5 to 1 m and was deposited under low energy, open marine, mid-ramp environments evidenced by its fine-grain size, nodular bedding, sparse fossil content, and association with other mid- and outer-ramp deposits (Figure 14.d). The skeletal wackestone subfacies ranges in thickness from 0.03 to 0.25 m and was deposited under low energy, moderately oxygenated, mid-ramp environments below fair weather wave base (Figure 14.c). This interpretation is based on its relationship with other deeper water facies (mudshale), occurrence of whole-fossil clasts, the shale interbeds, and the presence of bioturbation.
<table>
<thead>
<tr>
<th>Guling, northern India</th>
<th>Mikin Formation</th>
<th>Table 3</th>
</tr>
</thead>
<tbody>
<tr>
<td><strong>Subfacies</strong></td>
<td><strong>Bedding</strong></td>
<td><strong>Skeletal Components</strong></td>
</tr>
<tr>
<td>Nodular skeletal wackestone</td>
<td>Medium to thick (massive)</td>
<td>Abraded bivalves, gastropods, echinoderms, (conodonts)</td>
</tr>
<tr>
<td>Skeletal wackestone</td>
<td>Thin to medium</td>
<td>unabraded ammonoids, sparse bivalves, gastropods, conodonts</td>
</tr>
<tr>
<td>Skeletal mudstone</td>
<td>Medium</td>
<td>Sparse abraded bivalves, gastropods, echinoderms, (conodonts)</td>
</tr>
<tr>
<td>Skeletal mudstone-wackestone with mudshale interbeds</td>
<td>Medium</td>
<td>Sparse to rare fragments of bivalves, gastropods, echinoderms, (conodonts)</td>
</tr>
<tr>
<td>Dark mudshale interbedded with skeletal packstones</td>
<td>Thin</td>
<td>Shale: none Packstone: ammonoids, bivalves, gastropods, echinoderms (conodonts)</td>
</tr>
<tr>
<td>Dark mudshale</td>
<td>Thin</td>
<td>None observed</td>
</tr>
</tbody>
</table>
The four deep subtidal subfacies are *skeletal mudstone, skeletal mudstone-wackestone with mudshale interbeds, dark shale interbedded with skeletal packstones,* and *dark mudshale.* The *skeletal mudstone subfacies* (~1 m) subfacies was deposited in low energy, moderately oxygenated, outer-ramp environments evidenced by the occurrence of bioturbation, abundant mud, and sparse skeletal clasts (Figure 1.14a). The *skeletal wackestone with mudshale interbeds subfacies* ranges in thickness from 0.1 to 0.9 m and was deposited under low energy, poorly to moderately oxygenated, outer-ramp environments below fair weather wave base evidenced by the abundant shale partings, lack of bioturbation and sparse, abraded skeletal clasts, and its stratigraphic relationship with the deeper mudshale subfacies. The *dark mudshale interbedded with skeletal packstones subfacies* ranges in thickness from 0.05 to 2 m and was deposited above storm wave base in poorly oxygenated, outer-ramp environments (Figure 1.14b, c). This interpretation is based on the lack of fossils and sedimentary structures in the mudshale and the abraded, densely packed nature of the packstone lenses, and its facies relationship with the deepest water dark mudshale subfacies. The deepest subtidal *dark mudshale subfacies* (~1.5 m) was deposited under low energy, low oxygen outer-ramp environments below storm wave base evidenced by its fine-grain size, lack of fossils, bioturbation, and current-generated sedimentary structures, and its association with other interpreted deep subtidal subfacies.
Figure 1.14: Field photographs from the Guling, northern India section.  
a) In Guling, the base of the Smithian Mikin (skeletal wackestone subfacies (buff colored)) sits on a fault. Hiking pole (~1 m) for scale.  
b) The dark skeletal mudstone-wackestone with mudshale interbeds subfacies (pen for scale).  
c) Skeletal wackestone subfacies (buff) and the dark mudshale interbedded with skeletal packstones subfacies (person (~1.5 m) for scale).  
d) The nodular skeletal wackestone at the top of the section and with the dark mudshale interbedded with skeletal packstones (field book for scale).
1.4 Sequence Stratigraphy

1.4.1 Sequence Stratigraphy Terminology

Sequence stratigraphy is the study of genetically related sedimentary facies within the framework of chronostratigraphically significant surfaces. The fundamental unit of sequence stratigraphy is the sequence, which is defined as a relatively conformable, genetically related succession of strata bounded by unconformities or their correlative conformities (Mitchum, 1977). Parasequences and parasequence sets are the building blocks of sequences. Parasequences (or cycles) are defined as a relatively conformable, genetically related succession of beds or bedsets bounded by marine-flooding surfaces or their correlative surfaces and parasequence sets are parasequences that form a distinctive stacking pattern often bounded by major marine flooding surfaces and their correlative surfaces (Van Wagoner, 1988). Parasequences usually mimic the overall trend of the sequence, but on a shorter time scale.

Concepts associated with sequence stratigraphy were originally outlined using seismic stratigraphy and were based upon the vertical and lateral relationships of observed stratal geometries (Vail et al., 1977) and later evolved to include well log and outcrop studies (e.g. Van Wagoner et al., 1988; 1990). Since lateral facies trends and stratal geometries cannot be directly observed in the three outcrop field sites due to lack of seismic data and extensive lateral exposure, the traditional definitions and sequence stratigraphic concepts will be de-emphasized. Instead, the vertical stacking patterns and relationships will be used to infer lateral facies (i.e. Walther’s Law). The facies trends were used to infer depositional environments. The facies and associated subfacies were
ordered from shallowest to deepest (Tables 1-3). Based on the relative facies changes in the sections, a relative sea-level curve for each section was interpreted (Figure 1.15).

Systems tracts are defined as the linkage of contemporaneous depositional systems in a three-dimensional assemblage of lithofacies genetically linked by active (modern) or inferred (ancient) processes and environments (Fisher and McGowen, 1967; Brown and Fisher, 1977). The following will go through the original definitions of systems tracts, which will be used to describe and aid in the correlation of the three Smithian sections in the following sections. Once again, it is important to note that the descriptions used in this thesis will only loosely follow these definitions, as systems tracts terminology was intended to be used in cases extensive seismic data, lateral exposure, core and well logs.

The sequences are enveloped by sequence boundaries (SB), which are defined as significant erosional unconformities and their correlative conformities (Mitchum, 1977). These boundaries are the product of sea level falling faster than subsidence and sedimentation rate producing an unconformity or minimum accommodation (shallowest water sedimentation). The lowstand systems tract (LST) is the stratigraphically oldest systems tract and is represented by the sedimentary accumulation that straddles the lowest position of the relative sea level curve, often forming a prograding wedge at the base of a shelf margin (Van Wagoner, 1988). The transgressive systems tract (TST) follows the lowstand systems tract and is comprised of the deposits accumulated from the onset of the transgression until the time of maximum sea level rise rates (Van Wagoner, 1988). The TST represents a period of increasing accommodation space, which leads to a landward migration of facies and upward-deepening vertical facies patterns. The zone
that marks the time of maximum flooding or sea-level rise and separates the TST from the highstand systems tract (HST) is called the maximum flooding zone (MFZ) (Van Wagoner, 1988). The deposits that accumulate when sedimentation rates exceed accommodation rate increases constitute the HST (Van Wagoner et al., 1988; Jervey, 1988). The HST is characterized by shallowing-upward vertical stacking patterns.

1.4.2 Weber Canyon, northern Utah

The mixed carbonate-siliciclastic depositional sequence at Weber Canyon (~250 m) was measured and described on the north and south sides of the Weber River, adjacent to I-84 just west of Devil’s Slide. The underlying sequence boundary (SB) is characterized by the shallowest portion of fine-grained continental to marginal marine siliciclastics of the Woodside Shale (Smith, 1969), which contains mudcracks and lacks marine fossils. These observations suggest subaerial exposure and although, no evidence of an erosional surface was documented at Weber Canyon, the Woodside shale is interpreted to be the units deposited during a time of minimum accommodation. The Permian (Phosphoria Formation) and Triassic (Thaynes Formation) marine deposits bracket the Woodside Shale (Smith, 1969).
Figure 1.15: Smithian sequences from western North America (left and center; ~250 km apart) and northern India are correlated using sequence stratigraphy and biostratigraphy. The *Anasibirites* ammonoid zone has been identified by Stephen et al. (2010) for the Confusion Range section, Smith (1969) for the Weber Canyon section, and Krystyn (2004) for the Guling section. In addition, the *Meekoceras* ammonoid zone (Krystyn, 2004) is labeled for the Guling section and the *Neospatherodus waageni* conodont zone (Beaty, personal communication) for the Weber Canyon section to further constrain these sections to roughly Smithian time. The arrows on the Weber Canyon section indicate where each of the sketches in figure 1.16 are interpreted from.
The overlying transgressive systems tract (TST; ~150 m) is characterized by upward-deepening subfacies trends, including the silty lime mudstone, the abraded bivalve wackestone-packstone, the echinoderm grainstone, the silty skeletal lime mudstone. Superimposed upon these upward-deepening facies trends are ~1 to 3 meter-thick upward shallowing cycles. Biostratigraphy, although limited, provides insight into the relative time of deposition of the Lower Thaynes Formation. Beatty (personal communication) reports mid- to late Smithian *N. waageni* conodonts in the TST (~45 m). At the top of the TST (~153 m) there is an occurrence of the late Smithian *Anasibirities* ammonoid zone is documented (Smith, 1969).

The MFZ is composed of deep subtidal, dark calcareous mudshale (~15 m). The Sinbad Formation of central and eastern Utah represents the maximum landward extent of the marine limestones into the Thaynes seaway and it also contains the *Anasibirities* ammonoid zone (Lucas et al., 2007), suggesting that this interval represents the maximum flooding (Figure 1.3).

Upward-shallowing trends of the HST (~75 m) are represented by the following subfacies: lime mudstones interbedded with silty dark mudshale, the bivalve wackestone, and the silty lime mudstone. Superimposed in this upward-shallowing succession, are ~1 to 4 meter-scale upward-shallowing cycles.

The overlying SB is represented by the Decker Tongue (~20 m), which is composed of red siltstones and shales, which are interpreted as continental to restricted peritidal environments evidenced by occasional mudcracks and a lack of marine fossils (Smith, 1969). The Decker Tongue is the diagnostic unit used to separate the marine Lower Thaynes Formation from the Upper Thaynes Formation (Spathian).
Depositional Model for the Weber Canyon Platform

The Weber Canyon section is used as an example of how the stacking pattern for a single outcrop allows for lateral paleoenvironmental interpretations to be made across a basin. According to Walther’s Law, it is assumed that conformable vertically facies were deposited adjacent to one another in a basin. Using the four facies and seven subfacies defined above and their interpreted depositional environments to interpret relative (Figure 1.16).

Two of the three most prevalent and conformable subfacies at the base of the section, the silty lime mudstone and abraded bivalve wackestone-packstone are interpreted to have been deposited in a restricted shallow subtidal environment due to their relatively low energy, grain size, and abundance of mud. The third most abundant subfacies near the base is an echinoderm grainstone, which is interpreted to have been deposited in a high energy, open marine environment and is likely to have created a protective shoal barrier, so that the shallower facies listed above did not experience as much winnowing away of mud as would be expected for a typical shallow subtidal, open marine environment (Figure 1.16). Moving up section, the echinoderm grainstone is interbedded with a silty skeletal lime mudstone, which is interpreted as being deep subtidal based on the abundance of mud, parallel laminations, more abundant bioturbation at the base and increase in parallel laminations up section, indicating an overall decrease of oxygen and its relationship with the deepest subtidal facies, the dark calcareous mudshale up section. From these relationships, sea level is interpreted to be
Figure 1.16: Platform depositional model for the Weber Canyon section from Early TST to Late HST. There are seven subfacies that make up the Weber Canyon described on the right. The solid arrow corresponds to the location of Weber Canyon during deposition. These arrows are also marked on the stratigraphic column in figure 1.15.
rising, characterized by deeper water facies are depositing on top of shallower water facies.

Above the silty skeletal lime mudstone subfacies is the deepest subtidal subfacies, the dark calcareous mudshale, which represents the deepest water facies. The dark calcareous mudshale grades into the dark calcareous mudshale with interbedded lime mudstone facies, which is interpreted to have been deposited in the second deepest subtidal environment. This interpretation is based on the increased number of lime mudstone beds up section that contain sparse fossils, which indicated an environment of increasing oxygen and subsequent life, yet remains in a relatively low energy environment. This subfacies is also interbedded further up section with an intermediate subtidal, bivalve wackestone. This indicates that sea level is likely falling and the oxygen-poor deep subtidal facies are moving basin ward.

Further up section, the bivalve wackestone is interbedded with the same restricted shallow subtidal, low energy, devoid of fossils facies found at the base, which was interpreted to be the shallowest facies of the sequence. From these vertical relationships and some knowledge of other sections in the basin, it is evident that the Smithian at Weber Canyon represents a full sea-level transgression and regression (Figure 1.16).

1.4.3 Confusion Range, western Utah

The composite depositional sequence in the Confusion Range was measure and described at Cowboy Pass and Disappointment Hills, which are separated by ~16 km. Its total thickness (~240 m) is estimated from outcrop descriptions in this study combined with previous studies (Hose and Repenning, 1959; Collinson, 1976). The underlying SB
is represented by the unconformity between the Upper Permian Gerster Formation and the quartz-chert pebble conglomerate lying at the base of the Lower Thaynes Formation. Development of this unconformity has been attributed to uplift and subsequent erosion related to the Sonoma Orogeny (Collinson, 1976) or a drop in relative sea level (Newell, 1973). In either interpretation, at ~1 My of time is lost based on the fact that there have been no Griesbachian or Dienerian index fossils identified above the Permian Gerster Formation. The first diagnostic fossil occurs ~10 m above the unconformity and include lower Smithian Meekoceras ammonoid fossils (Stephen et al., 2010). Chemostratigraphy ($\delta^{13}$C), which is a tool used to correlate stratigraphic sections throughout the Phanerozoic, further supports that the Griesbachian and Dienerian stages are likely missing (Atudorei et al., 2007).

Deepening-upward characteristics of the overlying TST (~60 m) are characterized by the following subfacies: the clast-supported chert-quartz conglomerate, the vuggy (subtidal microbialite) lime mudstone-wackestone, the micro-gastropod-skeletal packstone-grainstone, the lime mudstone interbedded with calcareous mudshale rhythms, and the ammonoid-micro-gastropod wackestone. Smithian ages for the TST are provided by the Meekoceras, Inyoites, and Anasibirities ammonoid zones located in the ammonoid wackestone subfacies (Brayard et al., 2009; Stephen et al., 2010; Jenks, personal communication).

The MFZ is recognized by poorly exposed deep subtidal light (yellow) calcareous clayshale (~100 m thick) (Hose and Repenning; 1959). There have been no diagnostic fossils recorded in this stratigraphic unit; however, the first Anasibirities ammonoids are
recognized in the last recorded ammonoid-mirco-gastropod wackestone layer before the shale (Guex et al., 2010).

Upward-shallowing trends of the HST (~60 m) are represented by the bivalve wackestone subfacies overlain by the red silty limestone subfacies. Although there have been no age-diagnostic fossils identified, the first Spathian ammonoid bed is well above the marginal marine to continental red beds, which I interpret as the SBZ.

Marginal marine to continental red sandstones and shales of the Moenkopi Group (~75 m thick) define the upper SBZ (Guex et al., 2010). This unit is defined as the SBZ because it contains the shallowest marine deposits. This unit separates the marine Lower Thaynes Formation from the marine Upper Thaynes Formation (Spathian).

1.4.4 Guling, Himachal Pradesh, northern India

In northern India (Guling, Himachal Pradesh), the Smithian My-scale sequence (~14 m) is composed of mixed carbonate and siliciclastic TST (~5.5 m), siliciclastic dominant MFZ (~2 m), and a mixed carbonate and siliciclastic HST (~6.5 m) and was measured and described in the village of Guling, Himachal Pradesh (Appendix A). The underlying sequence boundary was not observed due to faulting, but at the Mud section ~10 km away it was characterized by nodular limestone subfacies (Krystyn et al., 2004), which represents the shallowest water facies for the Smithian and contains the Flemingities ammonoid zone (Krystyn et al., 2004; Galfetti et al., 2007). The upward deepening facies of the TST are characterized by nodular limestone and skeletal wackestone subfacies, overlain by dark shale. The Meekoceras ammonoid zone is found at the top of the TST. The maximum flooding zone is defined by the deepest water facies
and is composed of the deep subtidal (dark calcareous mudshales) (~1 m). The maximum flooding zone is above the *Meekoceras* ammonoid zone and below the *Anasibirities* ammonoid zone (Krystyn, 2004).

The upward shallowing facies trends of the highstand systems tract (HST) are characterized dark shale with abundant storm beds interbedded with skeletal wackestones and overlain by nodular limestone (~10 m). Twelve upward shallowing subtidal cycles occur in the HST characterized by dark shale with abundant storm beds (0.15 to 0.5 m thick) overlain by skeletal wackestone (0.08 to 0.3 m thick). Unlike Weber Canyon, the *Anasibirities* ammonoid zone is identified above the MFZ in the HST (Krystyn et al., 2004). Evidence of minimum accommodation (SB) is characterized by the nodular skeletal wackestone (Krystyn et al., 2004).

### 1.5 Global Correlation

Three Lower Triassic My-scale transgressive-regressive sequences are documented in the western United States (Embry et al., 1987; Paull and Paull, 1984), northern India (Embry, 1997; Krystyn et al., 2004), the Canadian Arctic (Haq, et al., 1987; Embry, 1997), northern Germany (Aigner and Bachmann, 1992; Embry, 1997), the northern Italian Alps (De Zanche et al., 1993; Ruffer and Zuhlke, 1995; Embry, 1997), Svalbard-Barents Shelf (northern Norway) (Van Veen et al., 1992; Embry, 1997), and eastern Siberia (Mørk et al., 1994; Embry, 1997). If these sequences can be correlated using biostratigraphy, then the sea-level changes generating the sequences could be eustatic in origin.
The Weber Canyon and Confusion Range are ~250 km away from each other and were deposited in the same basin. The TST for each section is confirmed to be Lower Thaynes based on existing ammonoid and conodont biostratigraphy (Stephen et al., 2010, etc.). For both locations, an occurrence of the Anasiberities ammonoid zone is directly below the MFZ. Further on shore, Anasirities is also documented in the Sinbad Formation, which is interpreted as the maximum extent of the Smithian transgression (Lucas et al., 2007). The sequence stratigraphic interpretation is based on the MFZ in Figure 1.15.

The TST at Weber Canyon has no known documentation of early Triassic biostratigraphy, which suggests that the conodonts and ammonoids were not living in this region during the early Triassic or that there is potentially time missing from the base of the WC section. The first identified age-diagnostic fossils are Middle to Upper Smithian in age and are located in the middle of the TST (N. waageni conodonts) and the top of the TST (Anasiberites). Since there is no known evidence of erosion, it is unlikely that there is missing time in the WC section. Instead, the water conditions (water depth, temperature, salinity, dissolved oxygen, etc.) were probably not favorable for the conodonts and ammonoids during the early Smithian.

The CR section, on the other hand, does have documentation of early, middle, and late Smithian ammonoids in the TST, suggesting that there is little Smithian time missing; however, it is likely that the G/D deposits were eroded away, since the Lower Thaynes Formation sits on top of the Permian Gerster Formation. Although, there are no age-diagnostic fossils in the HST, the first occurrence of an early Spathian Ammonoid (Tirolities) is stratigraphically above the interpreted SB. With
this in mind, it is likely that the HST is wholly or partially within the Upper Smithian, however, more data would be necessary to confirm this.

In the summer of 2009, I had the opportunity to measure and describe the Smithian section in Guling, India while on a research expedition with the intention of attempting a global correlation of the Smithian using the biostratigraphy outlined in Kristen et al. (2004; 2007). The Smithian portion of the Mikin Formation (~14 m) is much thinner than the Smithian of the western United States (~240 m) due to lower subsidence and sedimentation rates. The ammonoid and conodont bistratigraphy for the Guling, India section suggests that the entire transgressive-regressive sequence is within the Smithian.

The major difference between northern India and the western United States is that the Anasiberites ammonoid zone is located in the HST, rather than the TST and MFZ. Several scenarios could explain this. One situation is that the sea-level change, if in fact eustatic, did not happen all at once in the Smithian and, perhaps happened sooner in the Tethys Sea, which would mean that the sea-level change may not have been eustatic. Another solution is that the Anasiberities does continue through the interpreted HST for the western United States; however, there could be a preservation problem or the ocean conditions in the Panthlassa at that location were not conducive for the ammonoids to live there at that time. In either case, a more detailed study of the entire Lower Triassic is necessary to fully understand the global extent of the three transgressive-regressive sequences, which is beyond the scope of this thesis.
1.6 Conclusions

1. The Lower Triassic (Smithian) Lower Thaynes Formation in northern (Weber Canyon; ~250 m) and western (Confusion Range; ~240 m), Utah is composed of seven and eight subfacies, respectively. These subfacies can be grouped into four depositional facies representing restricted shallow subtidal, shallow subtidal, intermediate subtidal, and deep subtidal environments.

2. The Woodside Shale (marginal marine to continental) and the Decker Tongue (marginal marine) brackets the sequence at Weber Canyon. Mid- to late Smithian *N. waageni* conodont horizons occur within the TST and the Late Smithian *Anasibirites* ammonoid horizon at the top of the TST, which help to constrain the age of the sequence. The first Spathian (*Tirolities*) ammonoid horizon is found above the Decker Tongue.

3. An unconformity and conglomerate at the base of the Confusion Range section indicates that the section was subaerially exposed with at least 1 My of time missing between the Permian and Early Triassic. A very shallow marine red silty lime mudstone unit overlies the Lower Thaynes Formation at the Confusion Range. Early (*Meekoceras* and *Inyoities*) and Late (*Anasibirites*) ammonoid horizons have been identified in the TST. The first Spathian-diagnostic ammonoid horizon (*Tirolities*) is found above the restricted shallow marine red silty lime mudstone subfacies.
4. The Lower Triassic Smithian portion of the Mikin Formation (~14 m thick) in Guling, Himachal Pradesh, northern India is composed of two depositional facies and six subfacies that are grouped into representing intermediate and deep subtidal environments. Ammonoid biostratigraphy suggests that the entire transgressive-regressive sequence is within the Smithian. Unlike the western United States, the Anasibirities ammonoid zone is not found below the MFZ.

5. Using the maximum flooding zones defined for both the Confusion Range and Weber Canyon sections, as well as ammonoid and conodont biostratigraphy, it appears as though the two sections are roughly time equivalent.

6. The correlation between the two sections in the western United States was based on the maximum flooding zone and further constrained using the Anasibirities ammonoids and extended to the other side of the globe to a roughly time-equivalent India section based on the maximum flooding zone and ammonoid and conodont biostratigraphy. There correlation suggests that there could be a roughly time equivalent sea-level change between the three locations.
Chapter 2: Understanding the Early Triassic Climate using $\delta^{18}O$ of conodont apatite

Abstract

Samples were collected for oxygen isotopic analyses of conodont apatite from the two field locations (western North America) in a sequence stratigraphic framework to better understand the Smithian paleoclimate. Due to the an up to ~1.5‰ disparity of values between to two locations, additional samples were collected from Guling, northern India, the Sverdrup Basin in the Canadian Arctic, Wapiti Lake area in eastern British Columbia, Canada, and Bear Lake, southern Idaho. Conodont Alteration Index or CAI (a measure of thermal alteration based on color) was estimated for each of the locations listed above and range from 1.5 for Wapiti Lake, British Columbia and the Confusion Range to 5 for Guling, India. In addition, SEM images were taken to identify potential physical alteration of the conodonts for the Confusion Range (smooth surface with no signs of alteration), Weber Canyon (pitted surface with signs of potential alteration), and Guling (visibly the most pitted surface with signs of potential dissolution of the apatite).

The $\delta^{18}O$ values for the Weber Canyon range from ~14.4 to 15.8‰, Confusion Range range from ~16 to 16.9‰, Guling range from ~15.8 to 16.5‰, Wapiti Lake range from ~17.2 and 17.6‰, Sverdrup Basin range from ~14.5 and 14.8‰, and the Bear Lake value was ~16.5‰. Conodonts with a CAI of 3 or lower produced $\delta^{18}O$ values that most likely reflect the primary Smithian ocean isotopic values. From these values, the isotopic difference between tropical and subtropical latitudes is ~1‰. Assuming an ice-free ocean value of -1‰, sea-surface temperatures were calculated as 35 to 38 °C for the tropical and 32 to 34 °C for the subtropical regions. Both of these temperatures
indicate that the Smithian sea-surface temperatures were significantly warmer than the present ocean, and instead, most resemble the intense greenhouse sea-surface temperatures calculated for mid- to Late Cretaceous.

2.1 Introduction

Despite the range of multi-disciplinary research conducted to understand the potential causes and the recovery patterns of the Permo-Triassic extinction, little work has focused on establishing a detailed climate record during the ~5 My recovery interval after the extinction. Previous Early Triassic studies focus on the biotic recovery (Erwin 1996; Erwin 1996; Erwin and Hua-Zhang 1996; Hallam and Wignall 1997; Payne, Lehrmann et al. 2004; Payne, Lehrmann et al. 2004; Guex et al., 2010) and many of these studies cite climatic changes as a potential cause of the extinction and delayed biotic recovery; however, to date few studies have utilized $\delta^{18}O$ as a proxy for paleoenvironmental changes in this time interval.

$\delta^{18}O$ studies of marine minerals provide detailed insight into paleoclimate and paleoenvironmental studies including changes in ice volume, temperature, and seawater $\delta^{18}O$ composition (Urey et al., 1951; Longinelli, 1966; Kolodny et al., 1983; Luz et al., 1984; Kolodny and Luz 1991). $\delta^{18}O$ studies using biogenic phosphates are preferred over calcite because they are less prone to diagenetic alteration and non-equilibrium oxygen isotope fractionation has not been observed during biogenic apatite precipitation. Even though it is generally accepted that apatite is a good proxy for paleoclimate, there few published reports of how diagenesis affects primary isotopic values. The purpose of this
study is to 1) report the results from δ^{18}O apatite studies of Early Triassic (Smithian) marine successions in western North American and northern India, and 2) evaluate the reliability of conodont apatite δ^{18}O values in deep time (pre-Cenozoic) paleoclimate studies.

2.2 Geologic Background

Early Triassic paleogeography is dominated by the Pangaean supercontinent, which extended from ~85 °N to 90 °S and the globally extensive Panthalassa Ocean with its subtropical paleo-Tethys seaway extending into eastern Pangaea (Figure 2.1) (Ziegler et al., 1983; Blakey, 2010; Scotese, 2010). This study focuses on stratigraphic sections from the inland seaway in tropical, subtropical, and mid-latitudes along the western margin of Pangaea (western North America) and a southern subtropical passive margin within the southern Tethys Ocean on the eastern side of Pangaea (India) (Dubiel, 1994; Blakey, 2010; Scotese, 2010) (Figure 2.1). The locations for this study include Weber Canyon (northern Utah), Confusion Range (western Utah), Guling, Himachal Pradesh (northern Indian Himalayas), Wapiti Lake area (northeastern British Columbia, Canada), and the Sverdrup Basin (Canadian Arctic).

2.3 Early Triassic Climate

Previous studies of the Late Permian to Early Triassic suggest that the climate was uniformly warm with ice-free poles (e.g., Dickins, 1993, Frakes and Francis, 1988;
**Figure 2.1:** The supercontinent of Pangaea is fully assembled with the Panthalassa Ocean surrounding the eastern margin and the paleo- and neo-Tethys Seas bordering the eastern margin. Exposed land masses are dark gray and submerged continental platforms are light gray. Sites 1 and 2 are Weber Canyon, northern Utah and Confusion Range, western Utah, respectively, and were deposited in an intercratonic basin separated from the open Panthalassa Ocean by a volcanic arc. Site 3 is Guling, northern India deposited on a passive margin dips to the north towards the neo-Tethys Sea. Site 4 is Bear Lake, southern Idaho, site 5 is Wapiti Lake, northeastern British Columbia, and site 6 is the Sverdrup Basin, Canadian Arctic. Closed circles represent locations where I collected the samples and open circles are samples that were collected by Dr. Mike Orchard from the Canadian Geological Survey (modified from Blakey, 2010 and Scotese, 2010).
Evidence for this climate include: increased weathering rates determined from high latitude paleosols (Retallack, 1999; Michaelsen and Henderson, 2000), the migration of typically warm-water and warm-temperature flora and fauna to higher latitudes, such as calcareous algae in the Svalbard Basin (Wignall et al., 1998) and the appearance of cycads in Greenland (Looy et al., 2001). Particular ammonoid species have been found from tropical through mid-latitudes indicating only minor seawater temperature gradients (Brayard et al., 2006). No Early Triassic glacial deposits have been discovered suggesting polar to high-latitude regions did not have permanent ice or that the glacial deposits were not preserved (Frakes et al., 1992).

Several major negative shifts in δ\(^{13}\)C of up to 6 ‰ at the Permian-Triassic boundary and in the Early Triassic indicate major changes in the global carbon budget (Holser and Magaritz, 1987; Payne et al., 2004). The negative shifts are attributed to of the release of methane gas hydrates and subsequent rapid oxidation of methane to CO\(_2\) (Hallam and Wignall 1997). The increase in CH\(_4\) and CO\(_2\), increases the greenhouse affect and warms the climate (Berner, 1999; 2002).

In addition to geologic evidence of a globally warm climate, discussed above, general climate models across the Permo-Triassic boundary (Kidder and Worsley 2004) Kiehl and Shields, 2005), and geochemical modeling (Berner et al., 2002) suggests increase in the temperature at the boundary of 6-8 °C (Royer et al., 2004). This increase comes from an intensified greenhouse effect related to greenhouse gases in the atmosphere. Berner’s (2002) pCO\(_2\) models were conducted at 10 My time steps, which cannot detect shorter-scale climatic events and variability. In fact, the data available for
determining the Early Triassic paleoclimate is not abundant nor is it available for time scales of < 1 My. Even though there is abundant evidence to suggest that the Early Triassic was warm, Jefferson and Taylor (1983), Payne et al. (2004), Brayard et al. (2006), Preto et al. (2010) suggest that the climate was quite variable and much more complex than originally thought.

2.4 $\delta^{18}$O paleoclimate studies

Oxygen exists as three isotopes: $^{16}$O, $^{17}$O, and $^{18}$O with relative abundances of 99.759%, 0.037%, and 0.204%, respectively. Urey et al. (1951) pioneered the research efforts to fully understand the potential for oxygen isotopes to be applied as a paleothermometer. This is possible because $^{18}$O/$^{16}$O of calcite is dependent on a limited number of parameters: temperature and the composition of the ocean. $\delta^{18}$O, defined as:

$$\delta^{18}O = \left( \frac{\frac{^{18}O}{^{16}O}_{sample}}{\frac{^{18}O}{^{16}O}_{standard}} - 1 \right) \times 1000 \, ^{\circ}/oo$$

and can be measured in minerals containing oxygen, such as CaPO$_4$, H$_2$O, CaCO$_3$ (Craig, 1965). Emiliani (1955) was the first to apply this method of measurement and interpretation of the isotopic composition of oxygen to modern and ancient foraminifera, which set the stage for high-resolution paleoclimate studies.

One potential problem with using calcite is its susceptibility to diagenetic alteration in rocks older than the Late Mesozoic. Oxygen isotopic values in carbonates
can be altered during diagenesis by the addition of carbonate cement and the dissolution and re-precipitation processes. Early studies indicate that PO$_4^{3-}$ does not exchange oxygen with water during inorganic reactions (Tudge, 1960) and from there, later researchers began to test this method on fish bone apatite (Kolodny et al., 1983), phosphorite rocks (Shemesh et al., 1988), and conodonts (Luz et al. 1984) Conodonts are composed of carbonate fluorapatite and its composition is approximated by Ca$_5$Na$_{0.14}$(PO$_4$)$_3$$(CO$_3$)$_{0.016}$F$_{0.73}$(H$_2$O)$_{0.85}$ (Pietzner et al, 1968). Oxygen occurs in the apatite lattices in at least three sites; however, the method that we use isolates the PO$_4^{3-}$ group. Because apatite is less susceptible to diagenetic alteration than calcite, it is more likely to reflect the original temperature and composition of the ocean (Luz et al., 1984; Sharp et al., 2000). In addition, conodont apatite has been shown to produce reasonable paleotemperature estimates for the Silurian (Wenzel et al., 2000), Pennsylvanian (Joachimski, 2006) and Devonian (Joachimski et al., 2002; 2004; 2009; Elrick et al., 2009); therefore, we utilize conodonts as a proxy for deep-time paleoclimate reconstructions, much like foraminifera in the Cenozoic.

2.5 Methods

Three Early Triassic (Smithian) marine stratigraphic sections were measured and described (Chapter 1). Samples for $\delta^{18}$O analysis of conodont apatite were collected every 0.5 to 10 m from Weber Canyon and Confusion Range. Additional Smithian conodonts were obtained from Bear Lake, southern Idaho and from Wapiti Lake, northeastern British Columbia, the Sverdrup Basin in the Canadian Arctic (Canadian
samples were obtained from Dr. Mike Orchard at the Canadian Geological Survey), and Guling, northern India.

Bulk rock samples weighing between 10-25 kg were collected for each conodont sample. Samples were processed using standard preparation methods (Harris and Sweet, 1989); a density separation was performed using lithium metatungstate (LMT), and conodonts were handpicked under a binocular microscope. Conodont alteration index (CAI) values were estimated in order to determine the degree of thermal alteration based on the study by Epstein (1977). Representative examples were also mounted and imaged with a scanning electron microscope (SEM).

Conodont elements (0.3-1.2 mg) were converted to Ag₃PO₄ using a modified version of the methods described by O’Neil et al. (1994) and Bassett et al. (2007). The δ¹⁸O values were measured by reducing the Ag₃PO₄ to CO in a high-temperature conversion-elemental analyzer (TC/EA) connected to a Finnigan Mat 253 at the University of New Mexico, Albuquerque, New Mexico and New Mexico Tech, Socorro, New Mexico. Samples were run in duplicate, triplicate and occasionally quadruplicate to ensure reproducibility. All δ¹⁸O values were reported in ‰ relative to SMOW. The overall reproducibility was detected by replicate analyses of internal standards with an uncertainty of < 0.35 ‰ (1σ). The samples were normalized to the Tu-1 standard with a value of 21.11 ‰.
2.6 Results

2.6.1 CAI and SEM images of conodonts

Epstein (1977) developed the conodont alteration index (CAI), which measures the thermal alteration of the conodonts on a scale of increasing alteration from 1 to 8 based on the color of the conodont. Corresponding temperatures range from $< 50^\circ \text{C}$ for a CAI of 1 to $> 300^\circ \text{C}$ for a CAI of 5. The CAI estimates for each of the sections are shown in Table 2.1 and range from 1.5 to 4. SEM images of conodonts show varying degrees of physical alterations (Figure 2.2). Conodonts with a CAI of $\leq 2$ have a smooth and vitreous surface, whereas conodonts with a CAI of $\geq 3$ become increasingly more pitted and rough on the outer surfaces and on fractured surfaces.

<table>
<thead>
<tr>
<th>Location</th>
<th>CAI</th>
<th>Estimated temperature of alteration $^\circ \text{C}$</th>
</tr>
</thead>
<tbody>
<tr>
<td>Confusion Range, western Utah</td>
<td>1.5</td>
<td>50-90 $^\circ \text{C}$</td>
</tr>
<tr>
<td>Wapiti Lake, eastern British Columbia (Canada)</td>
<td>1.5</td>
<td>50-90 $^\circ \text{C}$</td>
</tr>
<tr>
<td>Bear Lake, southern Idaho</td>
<td>3</td>
<td>60-140 $^\circ \text{C}$</td>
</tr>
<tr>
<td>Weber Canyon, northern Utah</td>
<td>4</td>
<td>190-300 $^\circ \text{C}$</td>
</tr>
<tr>
<td>Sverdrup Basin, Canadian Arctic</td>
<td>4</td>
<td>190-300 $^\circ \text{C}$</td>
</tr>
<tr>
<td>Guling, Himachal Pradesh, northern India</td>
<td>5</td>
<td>$&gt; 300^\circ \text{C}$</td>
</tr>
</tbody>
</table>

Table 2.1
Figure 2.2: SEM images of typical conodont elements from a) the Confusion Range, western Utah, b) Weber Canyon, northern Utah, and c) Guling sections, northern India. Note the change in physical appearance of the elements, in particular, the Confusion range element is smooth, whereas the Weber Canyon and Guling elements are rough with possible evidence of dissolution.
2.6.2 $\delta^{18}O$ values

The measured $\delta^{18}O$ values range from 14.4 to 15.8 ‰ for Weber Canyon, from 16 to 16.9 ‰ for the Confusion Range, and 15.8 to 16.5 ‰ for Guling (Appendix B). The single Bear Lake sample has a value of 16.5 ‰ and two samples from Wapiti Lake have values of 17.2 and 17.6 ‰ and two samples from the Sverdrup Basin have values of 14.5 and 14.8 ‰ (Figure 2.3; Appendix B). When isotopic values are compared to the sequence stratigraphy at Weber Canyon or the Confusion Range, there are no systematic correlation, which would indicate a relationship between relative sea-level change and isotopic changes (Figure 2.4). Figure 2.6 illustrates the relationship between CAI and $\delta^{18}O$ values. With the exception of the two India samples, the observed trends show higher $\delta^{18}O$ values correspond to lower CAI values. This will be discussed further in the next section.

2.7 Discussion

2.7.1 $\delta^{18}O$ values

The measured Smithian $\delta^{18}O$ values range from ~14.4 to 17.6 ‰ across the western United States, the Canadian Arctic, and India. Of particular interest is that average $\delta^{18}O$ values between the two coeval Utah successions (Weber Canyon and the Confusion Range) record differences of up to 1.5 ‰ (Figure 2.3 and 2.4; Appendix B). This suggests that the seawater $\delta^{18}O$ values or seawater temperatures between the two localities, which
Figure 2.3: Averaged oxygen isotope measurements for the six sample locations arranged by paleolatitude. Approximate temperature calculations based on the Kolodny et al. (1983) equation are recorded on the top of the graph.
Figure 2.4: Oxygen isotopic values from the Confusion Range and Weber Canyon compared to measured and described stratigraphy. There are no apparent trends between the oxygen isotopes and the stratigraphy. In addition, there is an up to 1.5 per mil difference between the two locations. An interpreted sea-level curve determined in Chapter 1 is also plotted next to the stratigraphy derived from the sequence stratigraphic interpretations to the left of the stratigraphic columns.
are separated by ~300 km, were significantly different or that the conodont δ18O values at one or both of the Utah localities were diagenetically altered.

If the measured Utah δ18O values are primary, then the lower isotopic values at Weber Canyon may represent a significant influx of fluvial (meteoric) waters and a lowering of δ18O values of local coastal waters. The discharge of the modern Amazon River is one of the largest in the world; however, δ18O values in coastal waters near its mouth are depleted by only ~0.5‰ with respect to open-ocean Atlantic waters (Legrande and Schmidt, 2006). Given this comparison, it is not likely that the ~1.5 ‰ lower isotopic values recorded at Weber Canyon verses Confusion Range were caused by increased fluvial discharge because the amount of fresh water influx required to offset the Weber Canyon samples would require Early Triassic river systems several times larger than the modern Amazon. This amount of fluvial influx would preclude the occurrence of typical normal marine fossils found in the lower Thaynes Formation (i.e., conodonts, echinoderms, and ammonoids). The lower δ18O values at Weber Canyon might also represent higher seawater temperatures than at the Confusion Range. If this were the case, the ~1.5 ‰ difference between sites suggests greater than 4-6°C seawater temperatures at Weber Canyon than in the Confusion Range. Although possible, it is not likely such a drastic temperature changes would occur over the ~300 km distance between the two sites, which lay at a similar paleolatitude. In addition, these higher temperatures would likely prevent the occurrence of normal (though impoverished) marine fauna observed at Weber Canyon.
Alternatively, the higher δ¹⁸O values recorded in the Confusion Range might represent local seawater enrichment due to enhanced evaporation rates. Typical marine δ¹⁸O values in the enclosed and arid-climate Red Sea and eastern Mediterranean regions are enriched by up to 2‰ from adjacent open-marine waters in the upper 5 meters of the water column (Legrande and Schmidt, 2006). Such waters are typically too saline to support normal marine faunal types observed in the lower Thaynes Formation and suggests that the higher δ¹⁸O conodont values in the Confusion Range are not the result of evaporative seawater conditions.

The previous arguments suggest that the δ¹⁸O differences between the Smithian Utah successions are not result of differences in seawater chemistry or temperature. Instead, we suggest that δ¹⁸O differences are the result of diagenetic alteration, and in particular, the Weber Canyon conodonts record the effects of alteration related to elevated diagenetic temperatures. This interpretation is supported by the higher CAI values and SEM evidence of conodont recrystallization and/or pitting at Weber Canyon, while the Confusion Range conodonts have lower CAIs and more pristine textures under SEM (Figure 2.2). These interpretations are further supported by the CAI verses δ¹⁸O trends in Smithian conodonts from British Colombia and the Sverdrup Basin (Figure 2.5), where the conodonts with the lowest CAI values (least thermal alteration) record the highest δ¹⁸O values and most pristine SEM textures. In contrast to these trends, the northern Indian conodonts record the highest CAI values of 5, but their δ¹⁸O values lie within the same range as conodonts from the Confusion Range, British Columbia, and Idaho (CAI < 3). This suggests that CAI values alone cannot identify the potential effects of conodont alteration. Inorganic exchange between apatite phosphate and meteoric
Figure 2.5: Oxygen isotopic values versus conodont alteration index for the six sample locations. Conodonts with a CAI of less than or equal to 3 have higher oxygen isotopic values than those with a CAI of 4 or higher, with the exception of the Guling section.
waters would be very difficult at low temperatures (Tudge, 1960); however Blake et al. (1997) and Lecuyer et al. (1999) have found that early diagenetic microbial degradation of organic matter can stimulate secondary apatite precipitation into available voids resulting in a shift to lower oxygen isotope values. This may explain some of the observed Smithian trends.

Figure 2.6 shows the comparison between interpreted primary or near primary Smithian δ¹⁸O values (ranging from 15.85 to 17.6 ‰) from the Confusion Range, Idaho, British Columbia, and India to previously reported δ¹⁸O values from Ordovician through Triassic conodonts. The Smithian values are similar to those determined from the Early Ordovician greenhouse (Trotter et al., 2008) and overlap with lower isotopic values of greenhouse time intervals including the Middle Ordovician and Early and Late Devonian. Of particular interest is that the measured Smithian values are lower than limited data from the Earliest Triassic (Griesbachian-Dienerian; Korte et al., 2004) and the Middle and Late Triassic (Rigo et al., 2010). This suggests that Smithian seawater temperatures were warmer than the Earliest Triassic and warmer than the Middle and Late Triassic.

2.7.2 Paleoclimate interpretations

At present there is no routine independent proxy for temperature (like temperature-dependent species assemblages, TEX₈⁶, or alkenone paleothermometry) in deep geologic time as there is for the Cenozoic and latest Mesozoic, so estimates are made to decouple how much of the isotopic values are due to temperature verses seawater
Figure 2.6: All of the published oxygen isotopic values currently published from the Early Ordovician (Basset, 2007; Trotter et al., 2008), Middle and Late Ordovician (Trotter et al., 2008), Silurian (Wenzel et al., 2000), Early Devonian (Elrick et al., 2009; Joachimski et al., 2009), Middle Devonian (Joachimski et al., 2004; 2009; Elrick et al., 2009), Late Devonian (Joachimski and Buggisch, 2002; Joachimski et al., 2004; 2009), Middle Pennsylvanian (Elrick and Scott, 2010), Late Pennsylvanian (Joachimski et al., 2006), Early Triassic (Griesbachian/Dienerian) (Korte et al., 2004), and Middle to Late Triassic (Rigo et al., 2010) are plotted. The thickness of the line indicates approximately how many data points were published for each section. The thin tails on the ends of the Devonian values indicate <2 samples. The thinnest line indicates <10 samples, the medium thickness line indicates between 10 and 25 samples, and the thickest line (including this study) indicates that >25 samples were run. The pink shaded regions indicate warm periods in Earth history, the blue are cool, and the green is in between. Temperatures are estimated based on the Kolodny et al. (1983) equation assuming -1 $\delta_{sw}$ for warm periods, 0 for in between and +1 for cool periods. The range of temperatures are based on combining $\delta_{sw}$ to account for a the range of possible values between warm and cool times.
The average $\delta^{18}$O for globally mixed seawater ranges from -1‰ (ice free) to +1‰ (Pleistocene-size ice caps) with the present seawater value at ~0‰. If we assume the Confusion Range, British Columbia, Idaho, and northern India conodont isotope values record primary to near primary Smithian seawater values, then we can use the Kolodny et al. (1983) apatite temperature equation ($T(°C) = 113.3 - 4.38(\delta^{18}O_{phosphate} - \delta^{18}O_{seawater})$) to estimate seawater temperatures across the range of paleolatitudes.

Assuming ice-free conditions and a seawater isotope value of -1‰, seawater temperatures for the Confusion Range range from ~35 to 38°C, British Columbia range from ~32 to 34 °C, the single Idaho value is ~36 °C and India from ~37 to 40 °C. The fact the lowest paleolatitude sites (Confusion Range and Idaho ~10°N) record warmer temperatures than the higher paleolatitude site (British Columbian ~25°N) further support our interpretations that these samples represent primary to near primary isotopic values. These range of temperatures come from locations spanning at least 15° of tropical to subtropical paleolatitudes and indicate a gentle low-latitude to high-latitude temperature gradient during the Smithian. Although India had a higher paleolatitude at ~40 °S, the temperatures reported for this region are similar to those recorded in the subtropics and tropics in western North America. Transport of warm paleo-Tethys surface waters south to this region may explain these warm sea-surface temperatures (Figure 2.1). The dominance of platform carbonates (~40 m thick) spanning the Triassic in northern India supports these warmer temperature estimates.

For comparison, modern surface-water temperatures in the western tropical Atlantic range from ~27° to 29°C (Antonov et al., 1998), so our data suggest that the Smithian surface ocean was ~4° to 10°C warmer than the present. Estimates of surface
seawater temperatures during the mid to Late Cretaceous greenhouse range from ~31 to 36°C using TEX$_{86}$ (Forster et al., 2007; Bornemann et al., 2008) and ~30 to 37 °C using oxygen isotopes from well preserved planktonic foraminifera (e.g., Wilson et al., 2002; Bice et al., 2006; Forster et al., 2007; Bornemann et al., 2008). These Cretaceous tropical to subtropical sea surface temperatures are similar to the ranges estimated from the primary/near primary Smithian conodonts suggesting that Early Triassic climates were comparable to the well-studied Cretaceous greenhouse.

An alternative hypothesis to explain these unusually warm temperatures is that our assumption of -1‰ (ice-free) Smithian seawaters is incorrect and instead, Early Triassic seawaters had more depleted δ$^{18}$O values, perhaps as low as –3‰. Calculated seawater temperatures using such depleted δ$^{18}$O values would decrease and lie between ~25 to 30 °C, which are similar to modern tropical seawaters. The possibility of varying seawater δ$^{18}$O values over geologic time has been extensively discussed and argued (e.g., Veizer et al., 1997; 1999; Jaffres et al., 2007; and Shields et al., 2007); however, resolving this issue is beyond the scope of this research.

In summary, results from the studied stratigraphic sections recording primary to near primary oxygen marine isotope values suggests a gentle tropic-to-subtropical seawater temperature gradient and overall very warm seawater conditions during the Smithian which supports previous interpretations of greenhouse or hot-house climatic conditions in the Early Triassic (e.g. Dickins, 1993; Holser and Magaritz, 1987; Frakes and Francis, 1988; Frakes et al., 1992; Jefferson and Taylor, 1983; Wignall et al., 1998; Retallack, 1999; Berner, 1999; 2002; Michaelsen and Henderson, 2000; Looy et al., 2001; Spalletti et al., 2003; Kidder and Worsley, 2004; Payne et al., 2004; Royer et al.,
2004; Kiehl and Shields, 2005; Brayard et al., 2006; Preto et al., 2010). Estimated subtropical seawater temperatures from δ¹⁸O of conodont apatite (assuming -1‰ seawaters) immediately prior to the Permo-Triassic boundary range from 23 to 28 °C (Korte et al., 2004). Only two δ¹⁸O apatite data points are available for the Earliest Triassic (Griesbachian-Dienerian; Korte et al., 2004) and estimated temperatures from these two values suggest increasing seawater temperatures. Combining this sparse data with the primary to near primary Smithian δ¹⁸O apatite values from this study suggests continued warming conditions after the Permo-Triassic extinction and throughout the Early Triassic recovery interval. These unusually warm seawater temperatures may have aided in delaying the recovery after the Permo-Triassic extinction event.

2.7.3 What is driving My-scale sea-level change?

The original aim of this research was to collect samples of conodonts for δ¹⁸O values in a stratigraphic framework to understand what drove My-scale sea-level change. This is an especially important question for the Early Triassic because it appears to be uniformly warm and ice free. In chapter 1, we established that the My-scale sea-level change that occurred in the Smithian was likely eustatic. There is no apparent relationship between the δ¹⁸O values measured for the Weber Canyon and Confusion Range sections and paleoenvironmental change interpreted from facies changes (Figure 2.4). Because the Weber Canyon conodonts do not record primary δ¹⁸O values, they cannot be used to evaluate the relationships between My-scale sea-level change and climate. What is left unanswered is the origin of the Early Triassic My-scale sea-level
changes. Given the warm sea surface temperatures calculated for the tropical and sub-tropical climates, it is unlikely that sufficient glacial ice existed at high latitudes to drive glacial eustasy. This leaves the possibility that the Smithian sea-level changes were driven by thermal eustasy; however, more samples with a low CAI would be needed to test this hypothesis.

2.8 Conclusions

1. Smithian conodont samples with a CAI > 3 do not appear to record the primary $\delta^{18}$O values.

2. If we assume ice-free Smithian conditions, then tropical to subtropical sea-surface temperatures range from ~35-38 °C for the Confusion Range, western Utah (tropical) and ~31-34 °C for Wapiti Lake, northeastern British Columbia (subtropical). These temperatures are consistent with the Cretaceous greenhouse sea-surface temperature estimates for the similar latitudes and ~10 °C warmer than today’s tropical icehouse.

3. Regardless of the specific cause of the Permo-Triassic extinction, the seawater temperatures interpreted from this study suggest that very warm seawater temperatures influenced the protracted Early Triassic biotic recovery interval.
Appendices

Appendix A- Stratigraphic Columns ................................................................. 80
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APPENDIX A – STRATIGRAPHIC COLUMNS

The following pages contain stratigraphic columns for Weber Canyon, northern Utah, the Confusion Range, western Utah, and Guling, Himachal Pradesh, northern India. The Weber Canyon section was measured and described on two sides of the Weber River east of Ogden, Utah on I-84. The Confusion Range section was measured and described in three locations <40 km apart. The Guling section was measured and described in one location and observed in four other places in the same valley. All three locations represent Smithian marine transgressive-regressive cycles. The first page explains the symbols used and is followed by the data.
Appendix A: Legend

Sedimentologic/biologic structures:

- Symmetric Ripples
- Asymmetric Ripples
- Cross Bedding
- Laminations
- Bioturbation
- Argillaceous
- Calcareous
- Soft Sediment Deformation
- Vugs

Fossils:

- Crinoids
- Ammonoids
- Abraded bivalves
- Whole bivalves
- Gastropods
- Skeletal Fragments
- Peloids
- Sillicified Sponges
- Quartz sand

Bedding styles:

- Thick Bedding
- Medium Nodular Bedding
- Thin Bedding
- Shale

Conodont Sample (producing)

Conodont Sample (non-producing)

(t.s.) Thin section
Weber Canyon, northern Utah

Notes:

1 cm = 1 m
Weber Canyon, northern Utah

Notes:

TN-platy Nodular
dark, argilaceous
wk w/ poor exposure

Cleaner

t.s.

1 cm = 1 m
Weber Canyon, northern Utah

Notes:

1 cm = 1 m

Started measuring from the other side of the Weber River
Weber Canyon, northern Utah

Notes:

MFZ dark shale

1 cm = 1 m

Started measuring from the other side of the Weber River
Weber Canyon, northern Utah

Notes:

very little fizz at base

1 cm = 1 m
Weber Canyon, northern Utah

Notes:

Shaley interbeds

t.s.

t.s.

1 cm = 1 m
Weber Canyon, northern Utah

Notes:

Decker Tongue

- t.s.
- t.s.
- t.s.
- t.s.
- Reddish shale

1 cm = 1 m
Confusion Range, western Utah

Notes:

When in the field, I only measured section until here. One more same was collected right before the light calcareous shale subfacies at ~70 m. There is ~100 m of light shale between this section and the between top of the Lower Thaynes Formation, which is on the next page.

White
Beige (rose)
Gray
Gray
Reddish-brown

1 cm = 1 m
Confusion Range, western Utah

Notes:

Float/covered

1 cm = 1 m
Guling, northern India

Notes:

Storm beds

Shale interbeds

2 cm = 1 m
Guling, northern India

Notes:

Spathian Niti Limestone

SBZ

HST

M (S)  W (VFS)  P (FS)  G (MS)
APPENDIX B– OXYGEN ISOTOPES

The following pages contain oxygen isotope data for the six sections including the naming convention used and the measured stratigraphic location, if sampled in a stratigraphic framework. The bolded gray values represent already averaged values of up to 4 sample runs each, whereas the non-bold come from the final sample run and represent the raw data. CAI estimates are also shown in the right column.
<table>
<thead>
<tr>
<th>Location/sample name</th>
<th>Meters</th>
<th>$\delta^{18}$O values</th>
<th>Average</th>
<th>CAI</th>
</tr>
</thead>
<tbody>
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<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
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<td>11</td>
<td>14.78, 14.97</td>
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<td>4</td>
</tr>
<tr>
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<td>14.54</td>
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<td>4</td>
</tr>
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<td>218</td>
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<td>15.09</td>
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</tr>
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<td>Latitude</td>
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<td>Sverdrup Basin, Canadian Arctic</td>
<td>M1</td>
<td>15.02, 14.71</td>
<td>14.87</td>
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<td>M2</td>
<td>14.38, 14.70</td>
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<td>Wapiti Lake, eastern British Columbia</td>
<td>M3</td>
<td>17.86, 17.12</td>
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<td>M4</td>
<td>17.56, 17.09</td>
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